

- For electric dipole transitions, the spin-dependent part is

$$H_{int} = -\frac{e}{2mc} g_s \hbar \vec{S} \cdot \vec{\mathcal{H}}_{LM}^e$$

$$\vec{\mathcal{H}}_{LM}^e = \nabla \times \vec{A}_{LM}^e = \frac{1}{k} \nabla \times (\nabla \times \vec{A}_{LM}^m)$$

$$\begin{aligned} \text{But } \nabla \times (\nabla \times \vec{A}) &= -\nabla^2 \vec{A} + \nabla (\nabla \cdot \vec{A}) \\ &= +k^2 \vec{A} \end{aligned}$$

$$\therefore \vec{\mathcal{H}}_{LM}^e = k \vec{A}_{LM}^m$$

$$= +k \frac{i^L (2L+1)}{\sqrt{L(L+1)}} j_L(kr) \hat{L} C_{LM}$$

$$\approx +k \frac{i^L (2L+1)}{\sqrt{L(L+1)}} \frac{(kr)^L}{(2L+1)!!} \hat{L} C_{LM}$$

For magnetic multipole transitions, the orbital part is

$$H_{int}^m = -\frac{e}{mc} \vec{A}_{LM}^m \cdot \vec{p}$$

$$\vec{A}_{LM}^m \cdot \vec{p} = \frac{i^L (2L+1)}{\sqrt{L(L+1)}} j_L(kr) \hat{L} C_{LM} \cdot \frac{\hbar}{i} \nabla$$

$$\hat{L} C_{LM} \cdot \nabla = \frac{\hbar}{i} \underline{r} \times \nabla C_{LM} \cdot \nabla$$

$$= -\frac{\hbar}{i} \nabla C_{LM} \cdot \underline{r} \times \nabla$$

$$= -\nabla [j_L(kr) C_{LM}] \cdot \hat{L}$$

The spin part is

$$H_{int} = \frac{-e}{2mc} \hbar g_s \vec{S} \cdot \vec{\mathcal{H}}_{LM}^m$$

$$\vec{\mathcal{H}}_{LM}^m = \nabla \times \vec{A}_{LM}^m = k \vec{A}_{LM}^m$$

$$\propto \nabla(r^L C_{LM}) \quad \text{for } kr \ll 1.$$

Collecting the results, the total interaction operator is

$$H_g^{int}(\underline{k}, \underline{r}) = - \sum_{LM} \frac{(ik)^L}{(2L-1)!!} \sqrt{\frac{L+1}{2L}} \mathcal{D}_{Mg}^L(R)$$

$$\times \left\{ Q_{LM} + Q_{LM}' \cdot \frac{5}{L+1} - ig_s \mathcal{M}_{LM} \right\}$$

$$Q_{LM} = eg_L (r^L C_{LM})$$

$$Q_{LM}' = -k \mu_B g_s \mathcal{M}_{LM} (r^L C_{LM})$$

$$\mathcal{M}_{LM} = \mu_B \nabla(r^L C_{LM}) \cdot \left\{ 2g_L \frac{L}{L+1} + g_s \vec{S} \right\}$$

One must be very careful to include all the terms. For example, consider spin-forbidden electric dipole transitions - eg. $2^3P - 1^1S$.

$$H_{int} = -\frac{e}{2mc} \left[2 \underline{A} \cdot \underline{p} + \hbar \underline{\sigma} \cdot \underline{H} \right]$$

$$\underline{A}_{LM}^e = \frac{1}{k} \sqrt{\frac{L+1}{L}} \frac{(2L+1)}{(2L+1)!!} \nabla [(ikr)^L C_{LM}]$$

$$\underline{A}_{10}^e = -\frac{\sqrt{2}}{k} \nabla(z) = -\frac{\sqrt{2}}{k} \hat{k}$$

$$\begin{aligned} \underline{H}_{LM}^e &= \frac{+k i^L (2L+1)}{\sqrt{L(L+1)}} \frac{(kr)^L}{(2L+1)!!} \cos\theta \\ &= \frac{+k^2}{\sqrt{2}} (y \hat{i} - x \hat{j}) = \frac{+e\omega^2}{\sqrt{2}c^2} (y \hat{i} - x \hat{j}) \end{aligned}$$

The A.p term cannot connect directly the 2^3P and 1^1S states due to spin orthogonality, but it can connect the states indirectly through spin-orbit mixing.

$$\begin{aligned} |2^3P_1\rangle &= |2^3P_1\rangle_0 + \sum_{n \neq 1} \frac{|n^1P_1\rangle \langle n^1P_1| H_{so} |2^3P_1\rangle}{E(2^3P_1) - E(n^1P_1)} \\ &= |2^3P_1\rangle_0 + \sum_{n \neq 1} c_n |n^1P_1\rangle \end{aligned}$$

$$\text{Then } \langle 1^1S_0 | H_{int} | 2^3P_1 \rangle \quad c_n = \mathcal{O}(c^{-2})$$

$$\begin{aligned} &= -\frac{e}{2mc} \left[k \sum_n c_n \langle 1^1S_0 | 2 \underline{A} \cdot \underline{p} | n^1P_1 \rangle \right. \\ &\quad \left. + \hbar \langle 1^1S_0 | \underline{\sigma} \cdot \underline{H} | 2^3P_1 \rangle \right] \end{aligned}$$

$$= -\frac{e}{\hbar mc} \left[-\sqrt{2} \sum_n c_n \langle 1'S | p_z | n'P \rangle + \frac{\omega^2}{\sqrt{2} c^2} \langle 1'S | (y_i \hat{z} - x_j \hat{y}) \cdot \hat{\sigma} | 2^3P_1 \rangle \right]$$

In converting from the ~~length~~ velocity form to the ~~length~~ length form, we did not include the spin-orbit interaction in the Hamiltonian.

$$H_{so} = \frac{e \hbar}{4 m^2 c^2} \cdot \frac{Z e}{r^3} \frac{\hbar \cdot \hat{\sigma}}{\hbar} = \frac{\mu_B}{mc} \frac{Z e}{r^3} \frac{\hbar \cdot \hat{S}}{\hbar}$$

For ordinary spin-allowed transitions, we write

$$-\frac{e}{mc} \langle \alpha | p_z | \beta \rangle = \frac{e}{mc} \frac{m}{\hbar i} \langle \alpha | H_z - z H | \beta \rangle \\ = \frac{e}{i c} \omega_{\alpha\beta} \langle \alpha | z | \beta \rangle$$

Since we are now retaining terms of $\mathcal{O}(c^{-2})$, we must add to the above the spin-orbit commutator

$$\frac{e}{mc} \frac{m}{\hbar i} \langle \alpha | H_{so} z - z H_{so} | \beta \rangle \\ = \frac{e}{mc} \cdot \frac{e^2 Z^2}{4 m c^2} \langle \alpha | \frac{1}{r^3} \hat{\sigma} \cdot [\hat{L} z] | \beta \rangle \\ \downarrow \\ \frac{\hbar}{i} (y_i \hat{z} - x_j \hat{y})$$

It can be shown that this contribution exactly cancels the $\hat{\sigma} \cdot \hat{L}$ contribution.

Therefore, in the length form,

$$\langle 1^1S_0 | H_{int} | 2^3P_1 \rangle = -\frac{e}{2mc} \sum_n c_n \langle 1^1S_0 | z | n^1P_1 \rangle$$

.. all of this is much simpler in the Dirac theory. The Hamiltonian is simply

$$H = -e\phi + \beta mc^2 + \underline{\alpha} \cdot (c \underline{p} - e \underline{A})$$

At lowest order, the interaction operator is just

$$-e \underline{\alpha} \cdot \underline{A} \quad \alpha = \begin{pmatrix} 0 & \sigma \\ \sigma & 0 \end{pmatrix} \quad \beta = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$$

The above is not very useful because for many-electron atoms, it is actually the non-rel. wave function that is known best and so one must find the equivalent non-relativistic interaction operator.

Summary of Multipole Transition Probabilities.

1. Electric Multipole.

For spontaneous emission of a single photon from initial state ψ_i to final state ψ_f .

$$\hat{Q}_{LM}^{(e)} = \sqrt{\frac{4\pi}{2L+1}} r^L Y_L^M e r^* C_{LM}^* \quad \text{parity} = (-1)^L$$

$$A_{LM}^{(e)} = \frac{2(L+1)}{L(2L+1)[(2L-1)!!]^2} |\langle f | Q_{LM}^{(e)} | i \rangle|^2 \left(\frac{c\omega}{c}\right)^{2L+1}$$

$$\omega = E_i - E_f.$$

$$\tau = 2.4189 \times 10^{-17} \text{ sec.}$$

2. Magnetic multipole.

$$\vec{Q}_{LM}^{(m)} = \nabla(r^L C_{LM}^*) \left[\frac{e}{mc(L+1)} \vec{L} + \vec{\mu} \right] \quad \text{parity} = (-1)^L$$

$$\vec{\mu} = \frac{e\hbar}{2mc} \vec{\sigma}$$

$$A_{LM}^{(m)} = \frac{2(L+1)}{L(2L+1)[(2L-1)!!]^2} \left(\frac{\omega}{c} \right)^{2L+1} |\langle f | Q_{LM} | i \rangle|^2$$

Relative Magnitudes

$$\omega \sim Z^2$$

$$r \sim 1/Z$$

$$Q_{LM}^{(e)} \sim Z^{-L}$$

$$A_{LM}^{(e)} \sim \left(\frac{Z^2}{c} \right)^{2L+1} Z^{-2L} = \frac{Z^{2(L+1)}}{c^{2L+1}} \quad (n \neq n') = \frac{Z}{c^{2L+1}}$$

$$\text{For } E1, A_{1M}^{(e)} \sim \frac{Z^4}{c^3} \quad (n \neq n'), \quad (n=1)$$

- For spin-forbidden E1,

$$\frac{H_{50}}{\Delta E} \sim \frac{Z^4}{c^2 Z} = \frac{Z^3}{c^2}$$

$$A_{LM}^{e^2} \sim \frac{Z^{2(L+1)}}{c^{2L+1}} \cdot \left(\frac{Z^3}{c^2} \right)^2 = \frac{Z^{2(L+4)}}{c^{2L+5}} \quad (n \neq n')$$

$$A_{1M}^{(e)^2} \sim \frac{Z^{10}}{c^7}$$

$$\text{ratio} = \frac{7}{4} \cdot \frac{Z^8}{c^4} = (\alpha Z)^4 Z^5$$

$$\approx 1 \text{ at } Z = 9$$

$$Q_{LM}^{(m)} \sim \frac{Z^{-L+1}}{c} = \frac{Z^3}{c^{2L+3}} \quad (n=n')$$

$$A_{LM}^{(m)} \sim \left(\frac{Z^2}{c} \right)^{2L+1} \left(\frac{Z^{-2L+2}}{c^2} \right) = \frac{Z^{2(L+2)}}{c^{2L+3}} \quad (n \neq n')$$

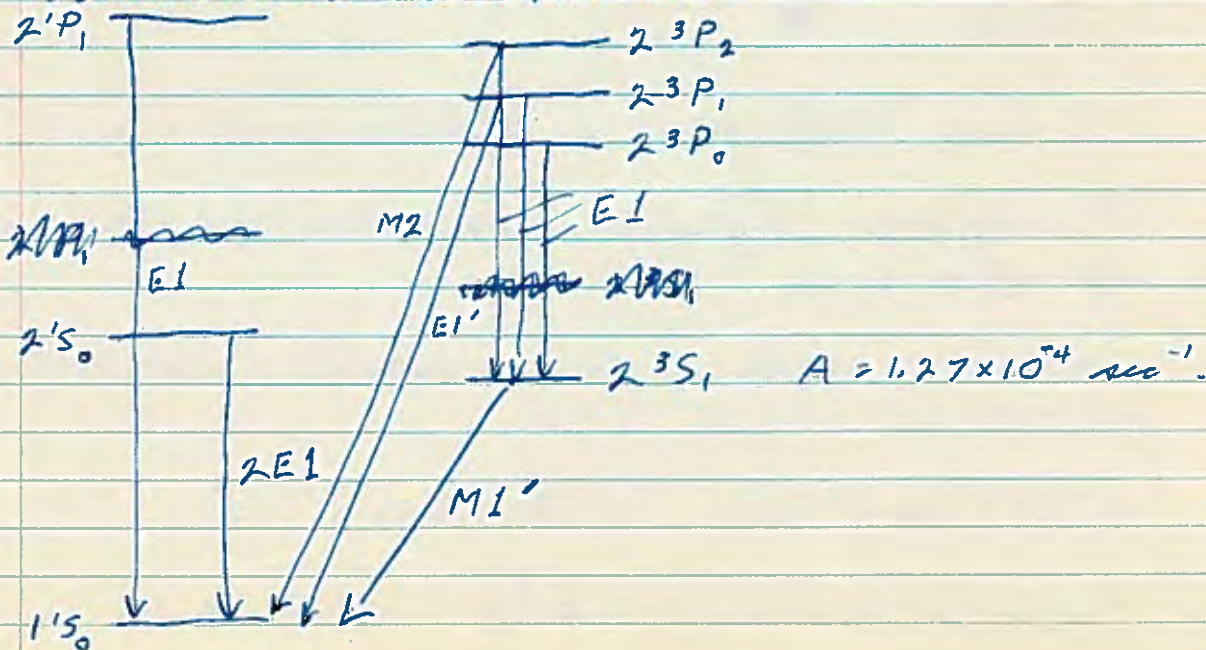
If the leading term of the Transition integral vanishes, as in $1s^2 \ ^3S_1 - 1s^2 \ ^1S_0$, the next correction term is $O(\omega^2 r^2 / c^2)$.

$$A_{LM}^{(m)'} \sim \frac{Z^{2(L+2)}}{c^{2L+3}} \left(\frac{Z^2}{c^2} \right)^2 = \frac{Z^{2(L+4)}}{c^{2L+7}} \sim \frac{A_{LM}^{(e)'}}{c^2}$$

The $O(c^{-2})$ correction terms to $\vec{Q}_{LM}^{(m)}$ have recently been calculated. The $1^1S - 2^3S$ transition integral is

$$\begin{aligned} & \langle 1s^2 \ ^1S_0 | Q_{10} | 1s^2 \ ^3S_1 \rangle \\ &= \mu_B \langle 1^1S | \sum_{j=1}^2 \left\{ \overbrace{L_{2j} + \sigma_{2j}}^{=0} - \frac{2}{3m^2 c^2} p_j^2 \sigma_{2j} \right. \\ & \quad \left. - \frac{1}{6} \left(\frac{\omega}{c} \right)^2 r_j^2 \sigma_j + \frac{Ze^2}{3mc^2} \frac{\sigma_{2j}}{r_j} \right\} | 2^3S_1 \rangle \end{aligned}$$

Transitions in Helium



For 2^3P_1

$$\frac{A(EI')}{A(EI)} \sim \frac{Z^{29}}{c^4}$$

$A(EI')$ becomes dominant for the ions beyond C II.

For 2^3P_2

$$\frac{A(M2)}{A(EI)} \sim \frac{Z^8/c^7}{Z^{10}/c^3} = \frac{Z^7}{c^4}$$

$A(M2)$ becomes dominant for the ions beyond C I XV.

The two-electron ions exist over a wider range of temp. and pressure conditions than any other ionization stage and thus are commonly observed.

Application to Planetary Nebulae

In the H II region, $T_e \sim 10^4$ K, $N_e \sim 10^4$ cm⁻³. Most atomic species are singly ionized.

$$\frac{N(He^+)}{N(He)} \sim \frac{10}{100}.$$

Time between inelastic collisions $\sim 10^3 - 10^4$ s.

Problem - explain observed He triplet spectrum.

Balance rates of formation and destruction of He 2^3S , 2^3P , 3^3D .

$$\beta(2^3S) N_e N(He^+) = N(2^3S) [A(2^3S - 1^1S)$$

$$+ g(2^3S - 2^1S) N_e + g(2^3S - 2^1P) N_e + R] \quad (1)$$

$$g(2^3S-2^3P)N(2^3S)N_e + \beta(2^3P)N_eN(He^+) \\ = N(2^3P)A(2^3P-2^3S) \quad (2)$$

$$g(2^3S-3^3D)N(2^3S)N_e + \beta(3^3D)N_eN(He^+) \\ = N(3^3D)A(3^3D-2^3P) \quad (3)$$

β = ^{effective} radiative recombination rate
 $He^+ + e^- \rightarrow He(2^3S)$

N_e = no. of electrons

g = ~~recombination~~ thermally averaged e^- collision induced transition rate.

R = photoionization rate.

The ratio $N = N(2^3S)/N(He^+)$ can be calculated directly from (1). Alternatively, we can solve (2) and (3) for N in terms of the observed intensity ratio

$$I = I(2^3P-2^3S)/I(3^3D-2^3P) = I(10830)/I(5876)$$

$$\frac{I(10830)}{I(5876)} = \frac{N(2^3P)A(2^3P-2^3S)}{N(3^3D)A(3^3D-2^3P)} \cdot \frac{5876}{10830} = \frac{1}{1.85}$$

$$N = \frac{1.85\beta(3^3D)I - \beta(2^3P)}{g(2^3S-2^3P) - 1.85g(2^3S-3^3D)I}$$

Oscillator Strength.

The oscillator strength for a transition between states n and n' is defined by

$$F_{nn'} = \frac{2m}{\hbar^2 e^2} \omega_{nn'} |\langle n | \hat{e} \cdot \vec{Q}_i^{(e)} | n' \rangle|^2$$

Then, in a.c.,

$$A_{nn'} = \frac{2\omega_{nn'}^2}{\hbar} F_{nn'} \tau^{-1}$$

The $F_{nn'}$ are dimensionless quantities satisfying the sum rule

$$\sum_{n'} F_{nn'} = N \quad N = \text{no. of electrons.}$$

For a transition between states labelled by $\gamma L M$, $\gamma' L' M'$,

$$\begin{aligned} F(\gamma L M \rightarrow \gamma' L' M') &= 2\omega |\langle \gamma' L' M' | \hat{e} \cdot \vec{Q}_i^{(e)} | \gamma L M \rangle|^2 \\ &= \frac{2\omega}{3} |\langle \gamma' L' M' | \vec{Q}_i^{(e)} | \gamma L M \rangle|^2 \\ &= \frac{2\omega}{3} |\langle \gamma' L' M' | \vec{r}_i | \gamma L M \rangle|^2 \end{aligned}$$

In terms of reduced matrix elements,

$$\langle \gamma' L' M' | \vec{r}_i | \gamma L M \rangle = (-1)^{L'-M'} \begin{pmatrix} L' & 1 & L \\ -M' & 0 & M \end{pmatrix} \langle \gamma' L' || \vec{r}_i || \gamma L \rangle$$

The averaged oscillator strength for the transition is

$$\begin{aligned} \bar{F}(\gamma L \rightarrow \gamma' L') &= \frac{1}{2L+1} \sum_{MM'} F(\gamma L M \rightarrow \gamma' L' M') \\ &= \frac{1}{2L+1} \frac{2\omega}{3} \sum_{MM'} \left| \begin{pmatrix} L' & 1 & L \\ -M' & 0 & M \end{pmatrix} \right|^2 \langle \gamma' L' || \vec{r}_i || \gamma L \rangle^2 \end{aligned}$$

In general

$$\sum_{M, M'} \left| \begin{pmatrix} L' & k & L \\ -M' & q & M \end{pmatrix} \right|^2 = \frac{1}{2k+1}$$

$$\sum_{0 \leq k} \frac{1}{2k+1} = 1$$

$$\therefore \bar{F}(Y_L \rightarrow Y'L') = \frac{2}{3} \frac{\omega}{(2L+1)} |\langle Y'L' | \sum_i \vec{r}_i | Y_L \rangle|^2$$

For the inverse process,

$$\bar{F}(Y'L' \rightarrow Y_L) = -\frac{(2L+1)}{(2L'+1)} \bar{F}(Y_L \rightarrow Y'L') \quad \text{go to 121.}$$

Quantum Mechanical Transition Probabilities

The Schrodinger equation is

$$H\Psi = i\hbar \frac{\partial}{\partial t} \Psi. \quad (1)$$

If H does not depend explicitly on time, then

$$\Psi_m(\vec{r}, t) = \psi_m(\vec{r}) e^{-iE_m t/\hbar}$$

$$H\psi_m(\vec{r}) = E_m \psi_m(\vec{r}).$$

H = total Hamilt.
Including both
atom + radiation
field.

Consider two states of the system ϕ_a and ϕ_b . These may not be eigenfns of H , but can be expanded in terms of them. The general sol'n. of (1) is

$$\Psi(\vec{r}, t) = \sum_m a_m \psi_m(\vec{r}) e^{-iE_m t/\hbar}$$

Suppose the state is initially in state ϕ_a

$$\phi_a = \Psi(\vec{r}, 0) = \sum_n a_n \psi_n \quad a_n = \int \psi_n^* \phi_a d\tau.$$

after time t , the state is

$$\begin{aligned} \Psi(\vec{r}, t) &= \sum_n a_n \psi_n e^{-iE_n t/\hbar} \\ &= \sum_n a_n e^{-iHt/\hbar} \psi_n \\ &= e^{-iHt/\hbar} \phi_a = e^{-iHt/\hbar} \Psi(\vec{r}, 0) \end{aligned}$$

The operator $e^{-iHt/\hbar} = 1 + \sum_{n=1}^{\infty} \frac{1}{n!} \left(\frac{-iHt}{\hbar} \right)^n$ is the time-development operator.

The quantity $\left| \int \phi_b^* \Psi(\vec{r}, t) d\tau \right|^2$

= prob. of finding state ϕ_b in time-developed state.

$$\therefore S_{a \rightarrow b}(t) = \left| \int \phi_b^* e^{-iHt/\hbar} \phi_a d\tau \right|^2.$$

Expansion of $S_{a \rightarrow b}$ in Eigenfunctions of H .

$$H|\alpha\rangle = E_\alpha|\alpha\rangle$$

$$H|\beta\rangle = E_\beta|\beta\rangle$$

$$\text{Expand } |\alpha\rangle = \sum_\alpha \langle\alpha|\alpha\rangle |\alpha\rangle$$

$$|\beta\rangle = \sum_\beta \langle\beta|\beta\rangle |\beta\rangle$$

$$S_{a \rightarrow b}(t) = \langle b | e^{-iHt/\hbar} | a \rangle \langle a | e^{+iHt/\hbar} | b \rangle$$

$$= \sum_{\alpha\beta} \langle b|\beta\rangle \langle\beta| e^{-iE_\alpha t/\hbar} |\alpha\rangle \langle\alpha| a\rangle$$

$$\times \langle\alpha|\alpha\rangle \langle\alpha| e^{+iE_\beta t/\hbar} |\beta\rangle \langle\beta| b\rangle$$

$$= \sum_{\alpha \beta} \langle b | \alpha \rangle \langle \alpha | a \rangle \langle a | \beta \rangle \langle \beta | b \rangle e^{-i(E_\alpha - E_\beta)t/\hbar}$$

Terms with $\alpha = \beta$ just give

$$\sum_{\alpha} \langle b | \alpha \rangle \langle \alpha | a \rangle \langle a | \alpha \rangle \langle \alpha | b \rangle = \sum_{\alpha} |\langle b | \alpha \rangle \langle \alpha | a \rangle|^2$$

For each term α, β , there is a corresponding term β, α .

$$S_{a \rightarrow b} = \sum_{\alpha} |\langle b | \alpha \rangle \langle \alpha | a \rangle|^2$$

$$+ \sum_{\alpha > \beta} \left\{ \langle b | \alpha \rangle \langle \alpha | a \rangle \langle a | \beta \rangle \langle \beta | b \rangle e^{-i\omega_{\alpha\beta}t} + \langle b | \beta \rangle \langle \beta | a \rangle \langle a | \alpha \rangle \langle \alpha | b \rangle e^{i\omega_{\alpha\beta}t} \right\}$$

Since $S_{a \rightarrow b}$ is real, the imaginary parts must cancel out, as they do. Also

$\langle b | \alpha \rangle \langle \alpha | a \rangle \langle a | \beta \rangle \langle \beta | b \rangle$ is real.

Since

$$|\langle b | \alpha \rangle|^2 = \langle b | a \rangle \langle a | b \rangle$$

$$= \sum_{\alpha \beta} \langle b | \alpha \rangle \langle \alpha | a \rangle \langle a | \beta \rangle \langle \beta | b \rangle$$

$$= \sum_{\alpha} |\langle b | \alpha \rangle \langle \alpha | a \rangle|^2 + 2 \sum_{\alpha > \beta} \langle b | \alpha \rangle \langle \alpha | a \rangle \langle a | \beta \rangle \langle \beta | b \rangle$$

the above becomes

$$S_{a \rightarrow b}(t) = -2 \sum_{\alpha > \beta} \langle b | \alpha \rangle \langle \alpha | a \rangle \langle a | \beta \rangle \langle \beta | b \rangle [1 - \cos(\omega_{\alpha\beta}t)]$$

The transition probability is obtained by taking the limit

$$P_{a \rightarrow b} = \lim_{t \rightarrow \infty} \frac{S_{a \rightarrow b}(t)}{t}$$

Using $\delta(x) = \frac{1}{\pi} \lim_{t \rightarrow \infty} \frac{1 - \cos(xt)}{x^2 t}$,

$$P_{a \rightarrow b}(t) = -2\pi \sum_{\alpha > \beta} \omega_{\alpha\beta}^2 \langle b | \alpha \rangle \langle \alpha | a \rangle \langle a | \beta \rangle \langle \beta | b \rangle \delta(\omega_{\alpha\beta})$$

The above is exact, but not very useful because the expansion coeffs. are not known.

The states $|a\rangle$ and $|b\rangle$ represent atomic wave functions in the absence of radiation field. They are expanded as l.c.'s of the exact $| \alpha \rangle$ and $| \beta \rangle$ which include the radiation field in H and therefore require a complete solution to Q.E.D.!

Perturbation Expansion:

- If a complete set of atomic eigenfns $|a\rangle, |b\rangle, |c\rangle, \dots$ are known, then the exact eigenfunctions ~~are found~~ $| \alpha \rangle, | \beta \rangle, \dots$ are found by diagonalizing $\langle H \rangle$.

Perturbation theory is useful if the off-diagonal matrix elements are not too large $O(1/c^2)$

$$| \alpha \rangle \approx | a \rangle - \sum_c \frac{\langle c | H | a \rangle}{\langle c | H | c \rangle - \langle a | H | a \rangle} | c \rangle$$

$$| \beta \rangle \approx | b \rangle - \sum_c \frac{\langle c | H | b \rangle}{\langle c | H | c \rangle - \langle b | H | b \rangle} | c \rangle$$

$$E_a \approx \langle a | H | a \rangle, \quad E_b \approx \langle b | H | b \rangle$$

If $\langle a | H | b \rangle \neq 0$, then

$$\langle a | \alpha \rangle \approx \langle b | \beta \rangle \approx 1$$

$$-\langle b | \alpha \rangle = \langle \beta | a \rangle = \frac{\langle b | H | a \rangle}{E_b - E_a}$$

Thus, in lowest order,

$$P_{a \rightarrow b} = \frac{2\pi}{\hbar^2} |\langle b | H | a \rangle|^2 \delta(\omega_{ab}).$$

If $\langle b | H | a \rangle = 0$, it is necessary to go to second order.

$$| \alpha \rangle = | a \rangle - \sum_{c \neq a} \frac{\langle c | H | a \rangle}{E_c - E_a} | c \rangle + \sum_{c, d \neq a} \frac{\langle d | H | c \rangle \langle c | H | a \rangle}{(E_c - E_a)(E_d - E_a)} | d \rangle$$

$$\langle a | \alpha \rangle = \langle b | \beta \rangle = 1$$

$$\langle \beta | a \rangle = \sum_c \frac{\langle b | H | c \rangle \langle c | H | a \rangle}{(E_a - E_b)(E_c - E_b)}$$

$$\langle b | \alpha \rangle = \sum_c \frac{\langle b | H | c \rangle \langle c | H | a \rangle}{(E_b - E_a)(E_c - E_a)}$$

$$\text{and } P_{a \rightarrow b} = \frac{2\pi}{\hbar^2} \left| \sum_c \frac{\langle b | H | c \rangle \langle c | H | a \rangle}{E_c - E_a} \right|^2 \delta(\omega_{ab})$$

$$E_c = \langle c | H | c \rangle \quad E_a = \langle a | H | a \rangle$$

The above is not completely correct because the initial and final states need not have exactly the same energy in second order - i.e. the transition is not completely sharp, but can have a finite width.

- Suppose $E_a \neq E_b$ i.e. $\langle a | H | a \rangle \neq \langle b | H | b \rangle$.

$$\text{Assume } |\langle a | H | a \rangle - \langle b | H | b \rangle| > |\langle a | H | a \rangle - \langle x | H | x \rangle|$$

$$\text{Assume a state } | x \rangle \text{ exists such that } |\langle a | H | a \rangle - \langle b | H | b \rangle| > |\langle a | H | a \rangle - \langle x | H | x \rangle|$$

$$\langle a | H | x \rangle \neq 0, \text{ but } \langle b | H | x \rangle = 0.$$

The exact eigenfunction $|\beta\rangle$ is expanded

$$|\beta\rangle = |\alpha\rangle - \sum_c \frac{\langle c|H|\alpha\rangle}{E_c - E_\alpha} |c\rangle$$

$$+ \sum_{c,d} \frac{\langle d|H|c\rangle \langle c|H|\alpha\rangle}{(E_\alpha - E_c)(E_c - E_\alpha)} |d\rangle$$

Then $\langle \alpha|\beta\rangle \approx \frac{\langle \alpha|H|\alpha\rangle}{E_\alpha - E_\alpha}$

$$\langle \beta|b\rangle \approx \sum_c \frac{\langle \alpha|H|c\rangle \langle c|H|b\rangle}{(E_\alpha - E_c)(E_c - E_b)}$$

Then, with $\langle \alpha|\alpha\rangle = \langle \beta|\beta\rangle = 1$,

$$P_{\alpha \rightarrow b} = \frac{2\pi}{\hbar^2} \sum_{x,c} \frac{\langle b|H|\alpha\rangle \langle \alpha|H|x\rangle \langle x|H|c\rangle \langle c|H|b\rangle \omega_{ax}}{(E_b - E_\alpha)(E_\alpha - E_x)(E_x - E_c)(E_c - E_b)} \times \delta(\omega_{ax})$$

The factor $\omega_{ax}^2 \delta(\omega_{ax})$ picks out $c = a$ from the sum. Then

$$\begin{aligned} P_{\alpha \rightarrow b} &= \frac{2\pi}{\hbar^2} \frac{|\langle b|H|\alpha\rangle|^2}{(E_\alpha - E_b)} \sum_x \frac{|\langle x|H|\alpha\rangle|^2}{(E_x - E_b)} \delta(\omega_{ax}) \\ &= \frac{2\pi}{\hbar^2} |\langle b|H|\alpha\rangle|^2 \cdot \frac{1}{\pi} \frac{\Gamma_a/2}{\omega_{ab}^2} \end{aligned}$$

where $\Gamma_a = \frac{2\pi}{\hbar^2} \sum_x |\langle x|H|\alpha\rangle|^2 \delta(\omega_{ax})$
= resonance width.